



Outer Solar System (OSS) A Planetary and Fundamental Physics Mission to Neptune, Triton and KBO

B. Christophe on behalf of OSS Team



r e t o u r s u r i n n o v a t i o n

OSS, new proposal for Cosmic Vision M3

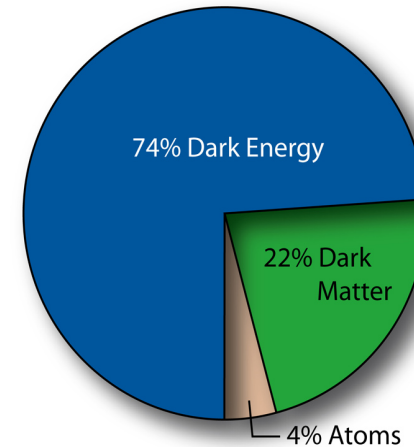
- In agreement with Fundamental Physics Roadmap Advisory Team of ESA (January 2010)
 - « Precise orbit determination of interplanetary spacecraft is an excellent tool to test general relativity and alternate theories of gravity [...] This is of particular interest for missions that cover large parts of the Solar system (e.g. missions to outer planets and Kuiper belt) in the light of scale dependent gravity. »
 - « A more modest M-class mission, although less sensitive and less complete, would be of strong interest for the study of large scale gravity, in particular when combined with planetary objectives, like for example the exploration of Neptune and its moon Triton and Kuiper belt objects. However such a mission might be difficult for ESA alone because of the lack of RTG power supplies in Europe [...] »

Fundamental physics scientific objectives

Objectives	Deep space gravity
Target precision	$\Delta a < 10 \text{ pm/s}^2$

Unification models predict deviations from General Relativity

“Dark matter” and “dark energy” are seen as gravitational anomalies; as long as they are not also observed through independent means, they may as well be interpreted as modifications of gravity laws at large scales

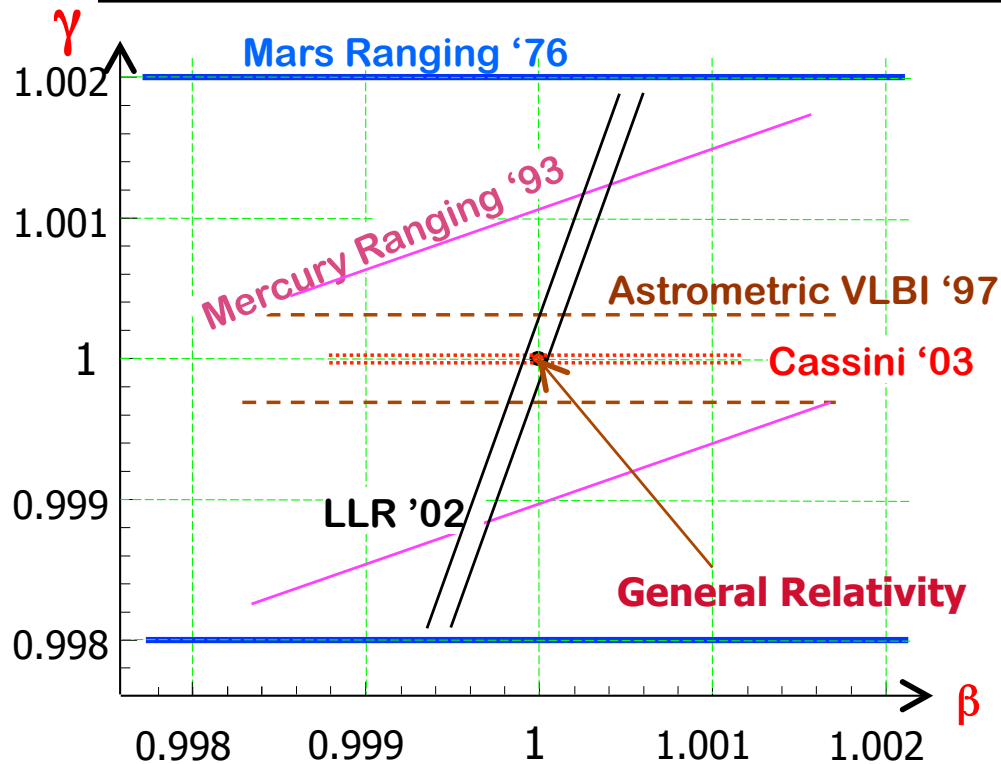


A new photo from NASA of the Bullet Cluster showing what is believed to be dark matter (represented in blue).
Source: NASA / CXC / CIA / STSci / Magellan / Univ. of Ariz. / ESO

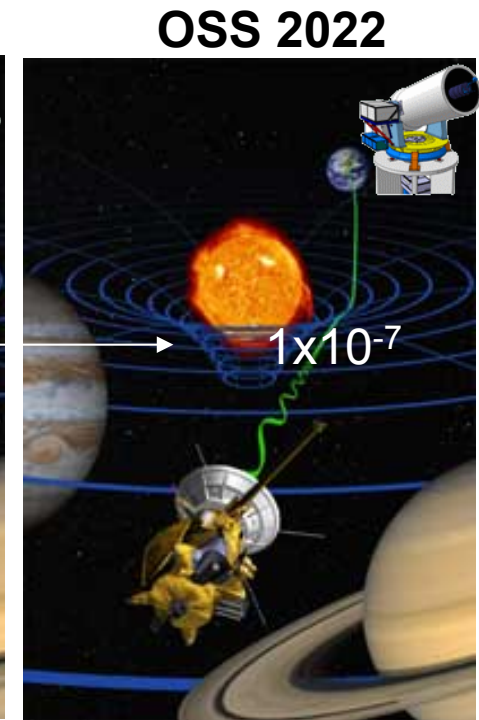
Importance of testing General Relativity at all possible scales beyond existing tests in the Solar System

Fundamental physics scientific objectives

Objectives	Deep space gravity	PPN parameter γ
Target precision	$\Delta a < 10 \text{ pm/s}^2$	$\Delta\gamma < 10^{-7}$



with Radio-Science



with Laser-Science

Courtesy: S. Turyshev

Planetary Objectives: Neptune, Triton & KBO

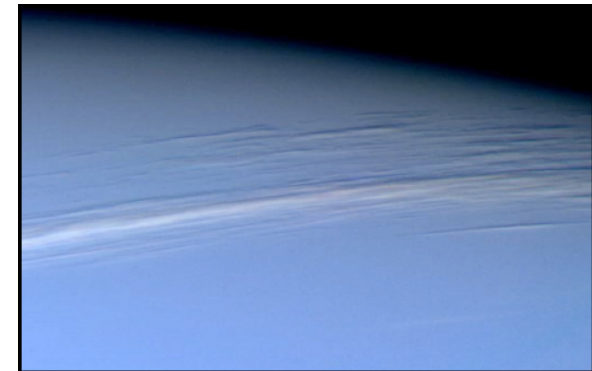
Last flyby of Neptune and Triton in 1989 by Voyager 2



Copyright, JPL/NASA

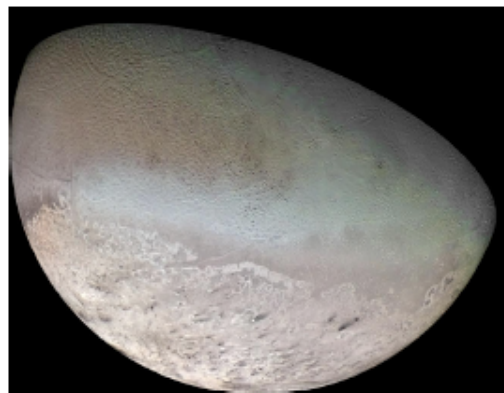
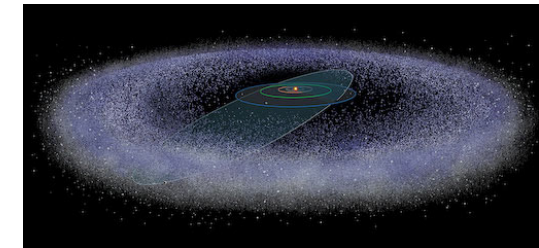
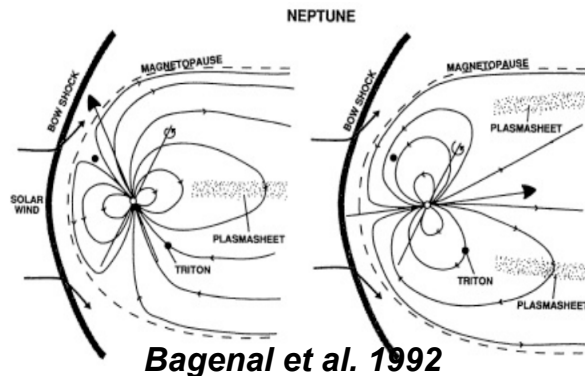
The ring system has changed since Voyager 2
The arcs evolved within 8 years

Neptune's wind possess the largest range in velocity in Solar System



Copyright, NASA

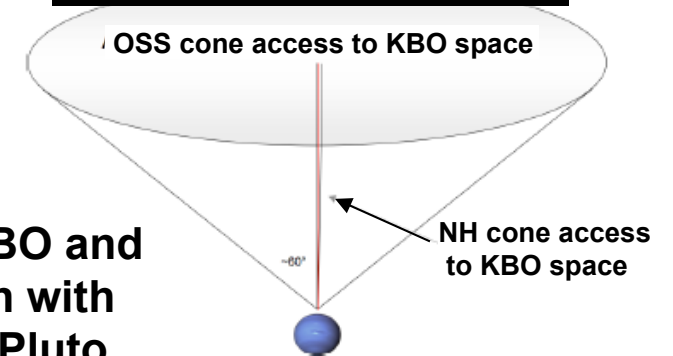
Neptune's magnetosphere is complex, with change over one Neptunian day at Voyager 2 epoch



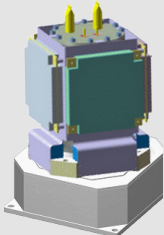
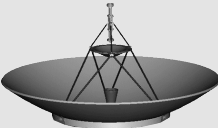
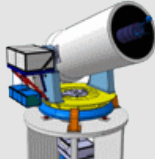
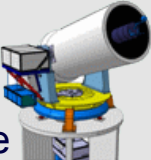
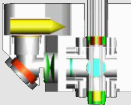
Copyright, NASA

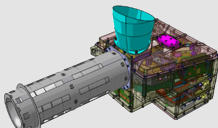
Triton has a young surface with very few impact craters
Triton could be a KBO

Flyby of a KBO and comparison with Triton and Pluto



OSS, Instrument suite

Instruments for Fundamental Physics	Mass Consumption
ACC DC Accelerometer 	3.5 kg 3.0 W
RSI Radio-Science 	3.0 kg 40.0 W
USO Ultra-Stable Oscillator	1.5 kg 5.5 W
LSI-1 Two-ways Laser Science 	25.0 kg 80.0 W
LSI-2 One-way Laser Science  +Clock Atomic Clock 	12.0 kg 12.0 W + 4.0 kg 15.0 W

Instruments for Planetary objectives	Mass Consumption
NIRS Near IR Spectrometer 	10.1 kg 7.5 W
UVS UV Spectrometer 	4.4 kg 4.4 W
NAC Narrow Angle Camera 	9.8 kg 14.0 W
TMI Thermal Map Imager 	3.4 kg 13.0 W
RPW Radio & Plasma Wave 	9.1 kg 5.9 W
MAG Magnetometer 	3.3 kg 3.0 W

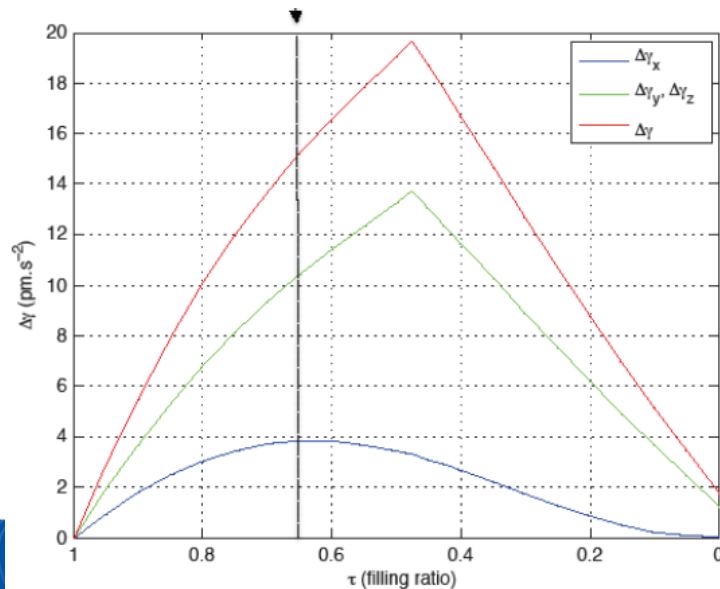
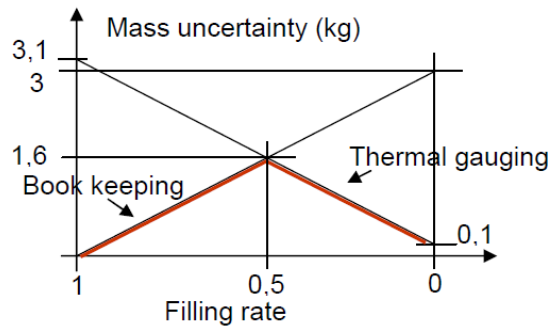
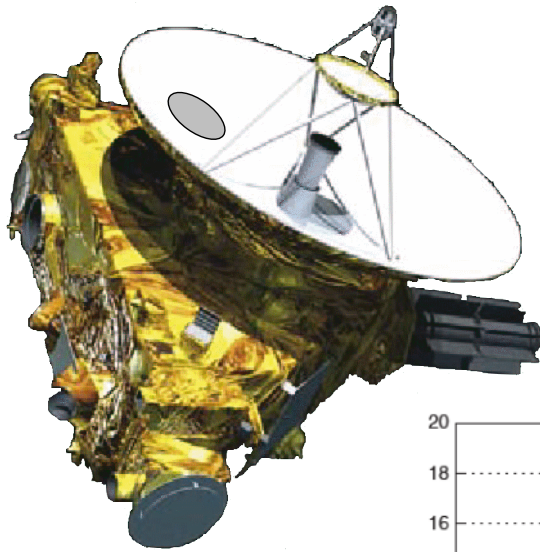
OSS, Measurement Scenario

Objectives	Instrument	Measurement Scenario
Deep space gravity	ACC, RSI	<p>The diagram illustrates the measurement scenario for various objectives. It features a central Sun (orange and yellow) with Earth (blue and white) and Neptune (blue and white) orbiting it. A Kuiper Belt Object (KBO) is shown as a small grey sphere in the distance. Labels include 'Kuiper Belt Object' with an arrow pointing to the KBO, 'Deep Space Gravity' at the bottom left, 'Eddington's parameter' near the Sun, 'Deep Space' near the KBO, and 'Neptune Triton' near Neptune. The background is a field of stars.</p>
PPN parameter γ	LSI, ACC, RSI	
Neptune Ring	NAC, NIRS, WAC	
Neptune Interior & Atmosphere	RSI, ACC, TMI, UVS, NAC, WAC	
Neptune Magnetic field	MAG, RPW, UVS, NIRS, ENA, PMS	
Triton	RSI, ACC, MAG, NAC, NIRS, UVS, TMI, PMS	
Kuiper Belt Object	RSI, ACC, MAG, NAC, NIRS, UVS, TMI, PMS	

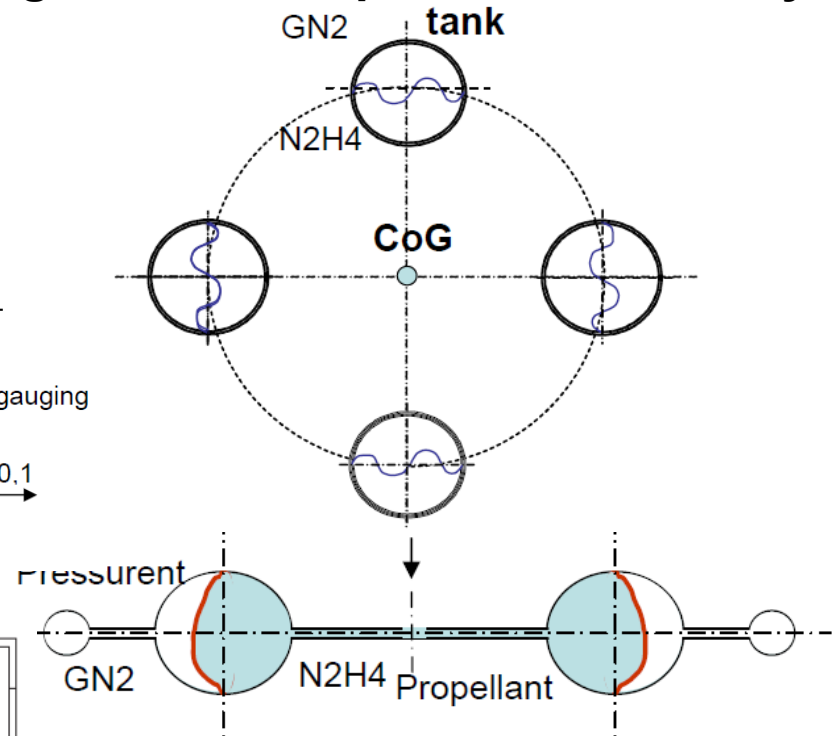
OSS, Spacecraft Design

Based on New Horizons,

- with HGA, 2,3 m diameter
- 2 MMRTG (125 W – 110 W)
- or 2 ASRG (146 W – 125 W)

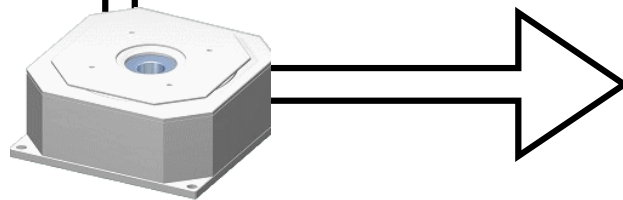
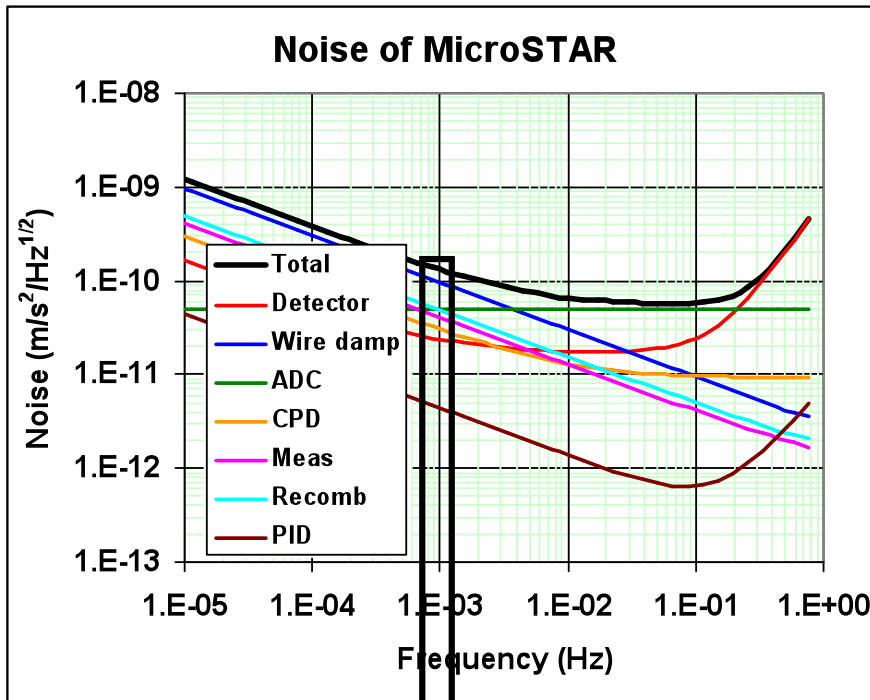


Management of Propellant Self-Gravity

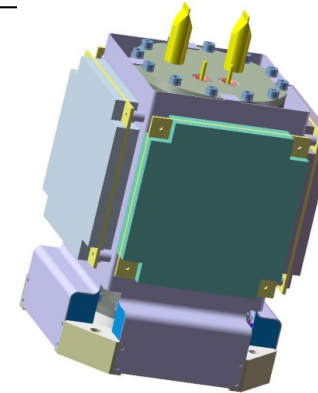


Self-gravity uncertainty of propellant during mission

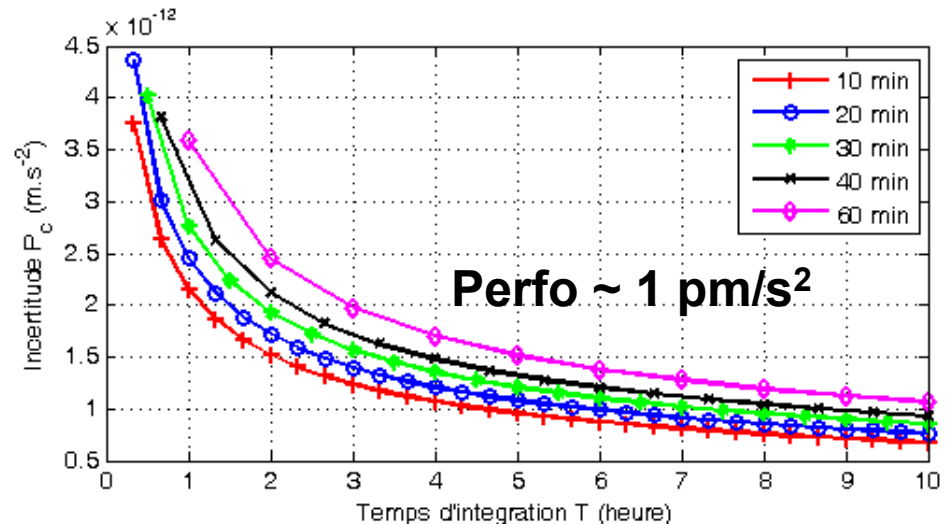
DC Accelerometer



Electrostatic accelerometer
 based on heritage of
CHAMP, GRACE and GOCE missions
 (developed by ONERA)

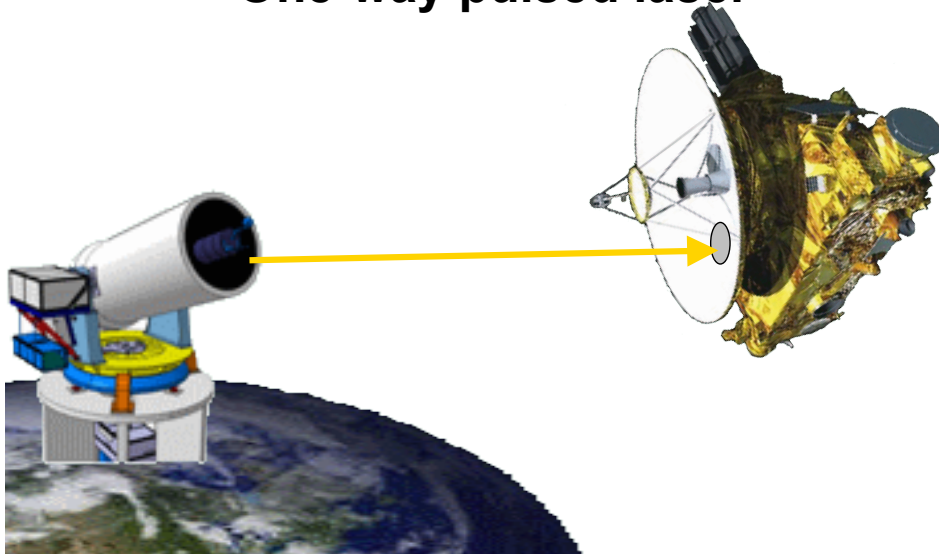


Bias Rejection System
 rotating stage with piezo-electric motor
 (developed by ZARM)

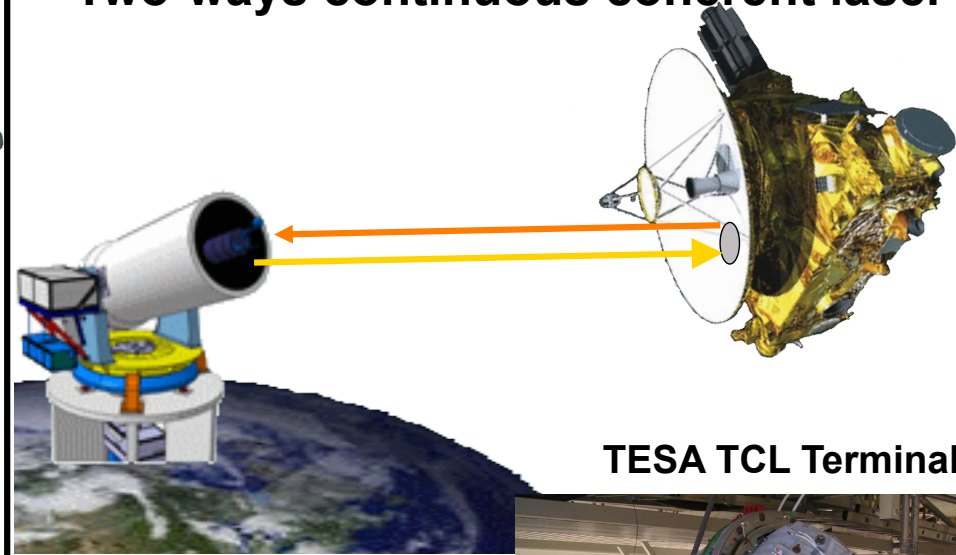


Laser Science: two concepts

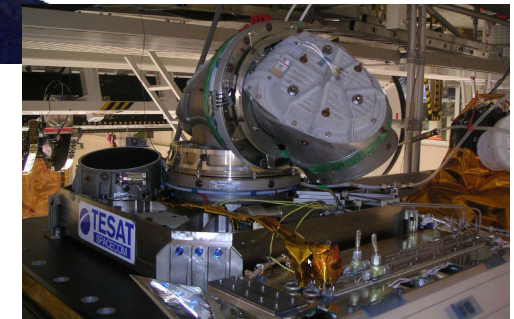
One-way pulsed laser



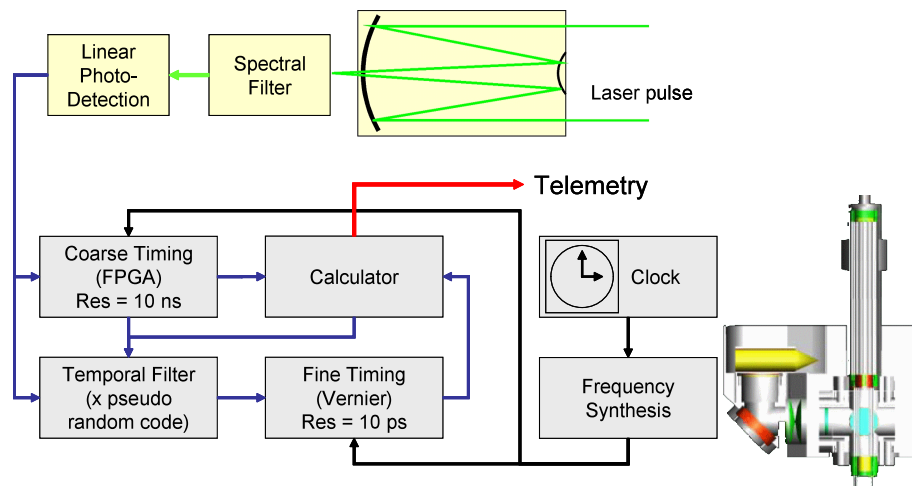
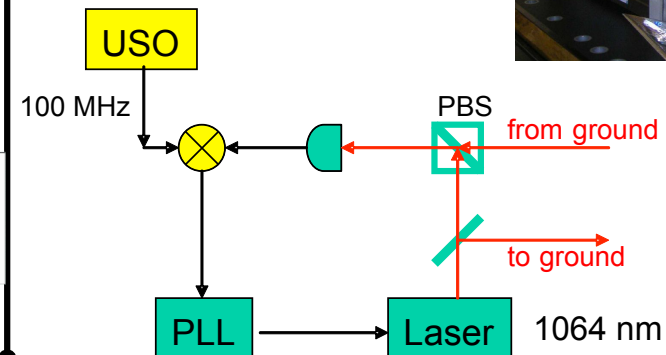
Two-ways continuous coherent laser



TESA TCL Terminal



Adaptative Optics on ground



Mercury Ion JPL or Horace Syrte clock

Conclusions

- **OSS, Fundamental Physics Mission with planetary objectives on Neptune/Triton/KBO**
 - In-line with ESA Fundamental Physics Roadmap
 - Proposed for next M3 call, in cooperation with US
 - Deep Space Gravity test constrained the S/C design (e.g. propellant for self-gravity)
 - DC Accelerometer accurate at 1 pm/s²
 - Laser-Science : 2 concepts to be analysed during the assessment phase

Acknowledgement to OSS scientific supporters

- **European side: PI B. Christophe (ONERA)**

- **CESR (France)** : *Dr. Nicolas André*
- **GeoAzur / OCA (France)** : *Dr. Gilles Métris, Dr. Etienne Samain*
- **IGEP, TU Braunschweig (Germany)**: *Prof. Karl-Heinz Glassmeier, Dr. Daniel Heyner*
- **Imperial College (United Kingdom)** : *Dr. Patrick Brown, Dr. Michelle Dougherty, Prof. Tim Sumner*
- **Institute for Geophysics and Meteorology / Univ. Köln (Germany)** : *Dr. Martin Pätzold*
- **Institut für Geowissenschaften / Max Planck Institute (Germany)**: *Dr. Frank Postberg*
- **Institute of Planetary Research / DLR (Germany)**: *Dr. Ralph Jaumann, Dr. Hauke Hussmann, Dr. Nicole Schmitz, Dr. Frank Sohl, Dr. Katrin Stephan*
- **Institute of Space System / DLR (Germany)**: *Prof. Hansjörg Dittus, Dr. Marco Sharringhausen, Dr. Tim van Zoest*
- **Instituto superior Técnico: Dr. Frederico Francisco, Dr. Paulo J.S. Gil, Dr. Jose Páramos**
- **JIVE (The Netherlands)** : *Dr. Sergei V. Progrebenko*
- **LESIA / Observatoire de Paris (France)**: *Dr. Baptiste Cecconi, Dr. Laurent Lamy*
- **LKB / ENS (France)** : *Prof. Jean-Michel Courty, Prof. Brahim Lamine, Prof. Serge Reynaud*
- **ONERA (France)** : *Bruno Christophe, Dr. Bernard Foulon, Benjamin Lenoir, Dr. Agnès Levy, Pierre Touboul*
- **Rutherford Appleton Laboratory (United Kingdom)**: *Dr. Robert Bingham*
- **SYRTE / Observatoire de Paris** : *Dr. Peter Wolf*
- **Universidade do Porto (Portugal)**: *Prof Orfeu Bertolami*
- **University of Oxford (UK)**: *Dr. Leigh Fletcher*
- **ZARM (Germany)** : *Dr. Claus Lämmerzahl, Dr. Hanns Selig*

- **US side: co-PI L. Spilker (JPL)**

- **Cornell University**: *Dr. Don Banfield*
- **JPL / NASA**: *Dr. John D. Anderson (retired), Dr. Sami W. Asmar, Dr. Glenn Orton, Dr. Kim R. Reh, Dr. Linda J. Spilker, Dr. Thomas R. Spilker*
- **Lowell Observatory**: *Dr. Will Grundy*
- **Planetary Science Institute**: *Dr. Candice Hansen*
- **UCLA**: *Dr. Johnatan Aurnou, Dr. Ravit Helled, Dr. Kunio Sanayagi*