



The Sagittarius tidal stream as a gravitationnal experiment in the Milky Way

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Dark matter or modified gravity ?

Begeman, et al. 1991 Sellwood & McGaugh 2005



► There is non visible matter → Adding a Dark Matter halo



Gravitational law is incomplete
Modifing the gravitational law



CMB - LSS

Works less well

ESA, the Planck Collaboration & SDSS



Works well

- CMB LSS
- Galaxy clusters

Works less well

Bullet Cluster



Works well

CMB - LSS

Galaxy clusters

Works less well

Baryonic Tully-Fisher Relation $\log M_b = 4 \log (V_f) - \log (G a_0)$



Famaey & McGaugh, 2012



CMB - LSS

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Works less well

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Baryonic Tully-Fisher Relation

 $\log M_b = 4 \log (V_f) - \log (G a_0)$

Mass discrepancy – acceleration relation

MOdified Newtonian Dynamics

Proposed in 1983 by Mordelai Milgrom

$$g = g_N$$
 if $g >> a_0$
 $g = (g_N a_0)^{1/2}$ if $g << a_0$

QuMOND : Quasi – linear formulation of MOND

Mondian Poisson equation :

$$\nabla^2 \Phi = 4 \pi G(\rho_b + \rho_{ph})$$

Phantom dark matter :

$$\rho_{ph} = \nabla \left[\widetilde{v} \left(\frac{|\nabla \Phi_N|}{a_0} \right) \cdot \nabla \Phi_N \right]$$

Relativistic formulations exist : BIMOND, RAQUAL, TeVeS ... but problem with the CMB and clusters

MOdified Newtonian Dynamics

Interpolation functions :

$$\widetilde{\mathbf{v}}(y) = \frac{(1+4y^{-1})^{(1/2)} - 1}{2}$$

Simple

$$\widetilde{v}(y) = \left[\frac{1 + \sqrt{1 + 4y^{-2}}}{2}\right]^{1/2} - 1$$

> Exponantial

$$\widetilde{v}(y) = (1 - e^{-y^2})^{-1/4} + \frac{3e^{-y^2}}{4} - 1$$



Success of MOND



Rotation curves of spirals galaxies well reproduced

Also for LSB galaxies

Explain the galaxy Gravitational Lensing



Milgrom et al, 2013

Famaey & McGaugh, 2012

New test of MOND : Streams

- Stream formed by merger of a dSph with a host galaxy
- Stream are important to trace the shape of the potential outside of the plane
- No dynamical friction in MOND
- Role of the External Field Effect (EFE)

Sagittarius Dwarf Spheroidal (dSph Sgr)

- Discovered in 1994 by R. IBATA
- $M_{\star} = 2.1 \ 10^7 M_{\odot}$
- r_h = 2.587 kpc
- D_{sun} ~ 25 kpc
- Long stream well visible



Koposov et al, 2012

New test of MOND : Streams

Milky Way model

- Disk (Thin, thick, ISM)
- Bulge
- Without DM halo
- $M_{MW} = 5.64 \ 10^9 M_{\odot}$

Sgr dSph model

- King profile
- $M_{Sgr,dSph}(t=0 \text{ Gyr}) = 4.67 \ 10^7 \text{ M}_{\odot}$
- r, (t= 0 Gyr) = 0.84 kpc



New test of MOND : Streams



Conclusion





Compare generically the simulations and the currents observations

- Simulate other streams in the Milky Way and in other galaxies (M 31, ...)
- Make predictions and compare them with the future Gaia results



