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Actual situation of nutation : Comparison between official VLBI solutions

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 Observatoire
de Paris | SYRTE
Systèmes de Référence Temps-Espace

Introduction

FIELD : Geodetic Very Long Baseline Interferometry

Purpose of the thesis :

Astrometric :

Situation :

ICRF2 (soon ICRF3) Cf « The second Realization of the International Celestial Reference Frame by Very Long Baseline Interferometry », IERS technical note no.35

Geodesic :

Precession-nutation

Earth Rotation

ITRF

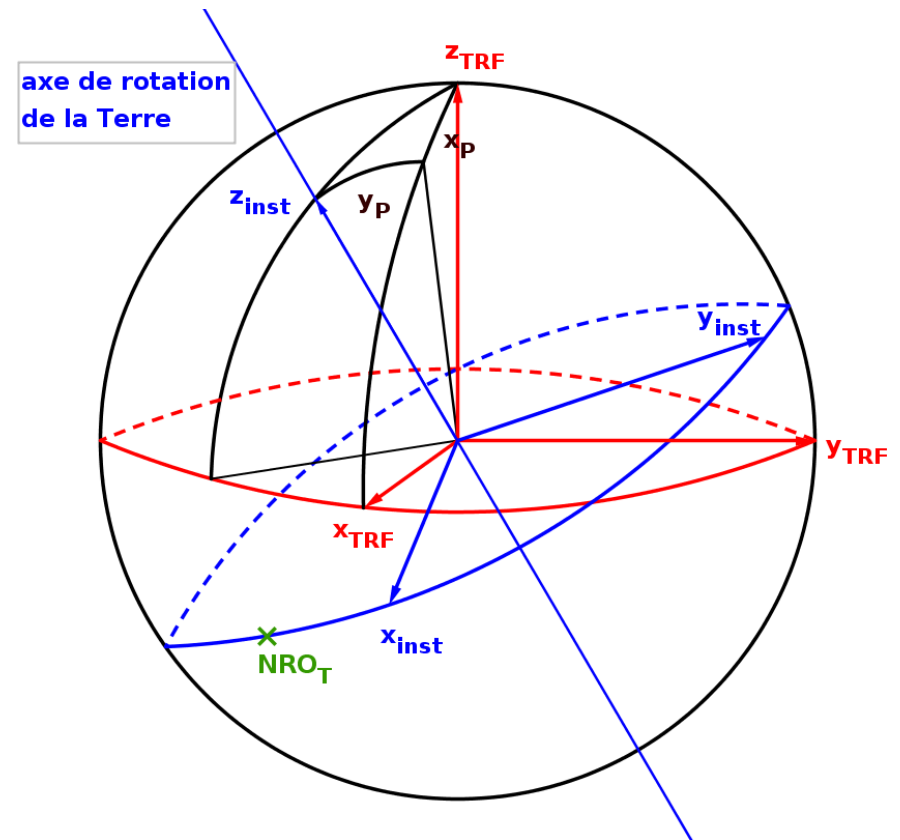
Some basic knowledge

$$[\mathbf{CRF}] = \mathbf{PN} \cdot \mathbf{A} \cdot \mathbf{PM} [\mathbf{TRF}]$$

Where \mathbf{PN} , \mathbf{A} and \mathbf{PM} are rotation matrix products.

\mathbf{PM} allows to change :

- From the terrestrial reference frame
- To the immediate frame of the date
(having as pole the Earth rotation pole and possessing a non-rotating origin according to the Earth crust)



⇒ **Earth Polhodie**

→ Parameter : x_p and y_p

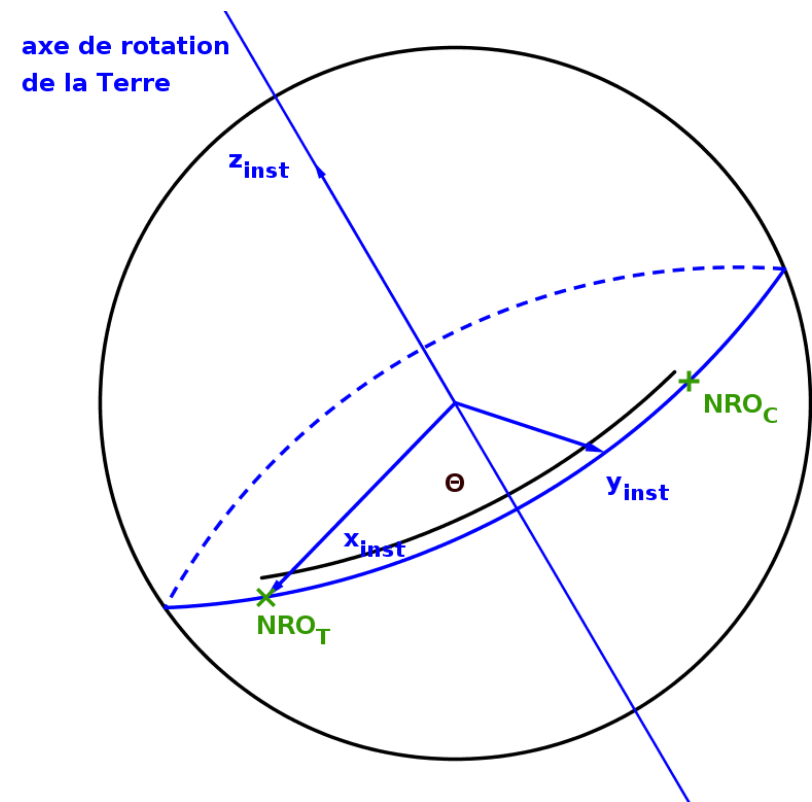
Some basic knowledge

$$[\mathbf{CRF}] = \mathbf{PN} \cdot \mathbf{A} \cdot \mathbf{PM} [\mathbf{TRF}]$$

Where \mathbf{PN} , \mathbf{A} and \mathbf{PM} are rotation matrix products.

\mathbf{A} allows to change :

- From the immediate frame of the date
(having as pole the Earth rotation pole)
 - Non-rotating origin according to the Earth crust
- To the immediate frame of the date
(having as pole the Earth rotation pole)
 - Non-rotating origin according to the celestial sphere



⇒ **Earth rotation angle**

→ Parameter : **UT1**

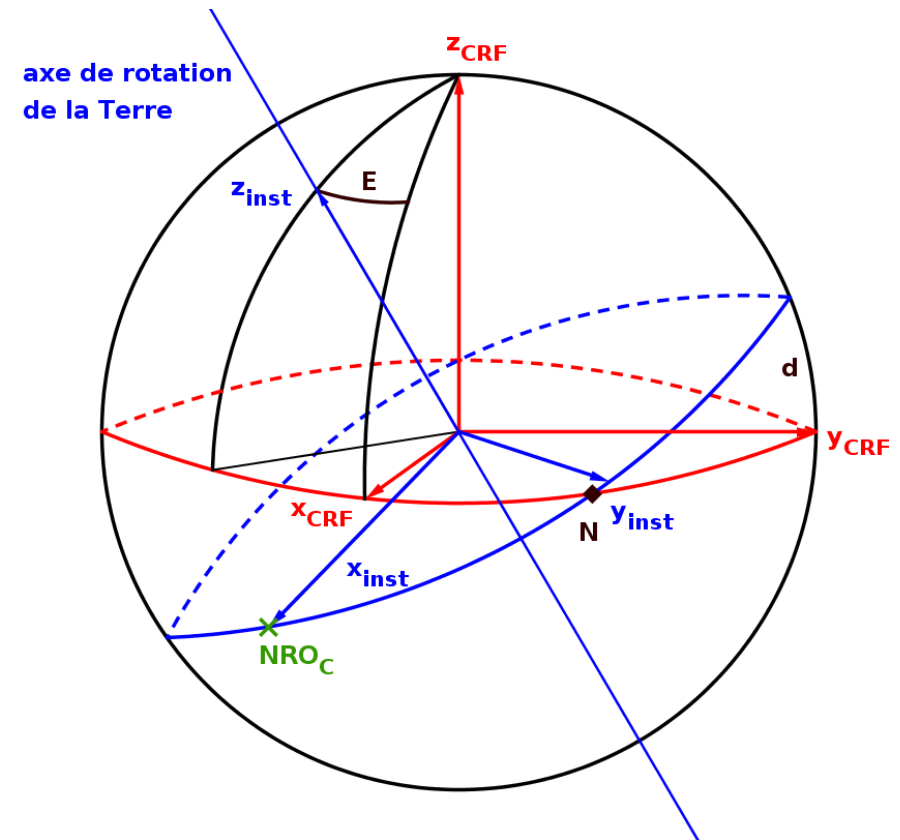
Some basic knowledge

$$[\mathbf{CRF}] = \mathbf{PN} \cdot \mathbf{A} \cdot \mathbf{PM} [\mathbf{TRF}]$$

Where \mathbf{PN} , \mathbf{A} and \mathbf{PM} are rotation matrix products.

\mathbf{PN} allows to change :

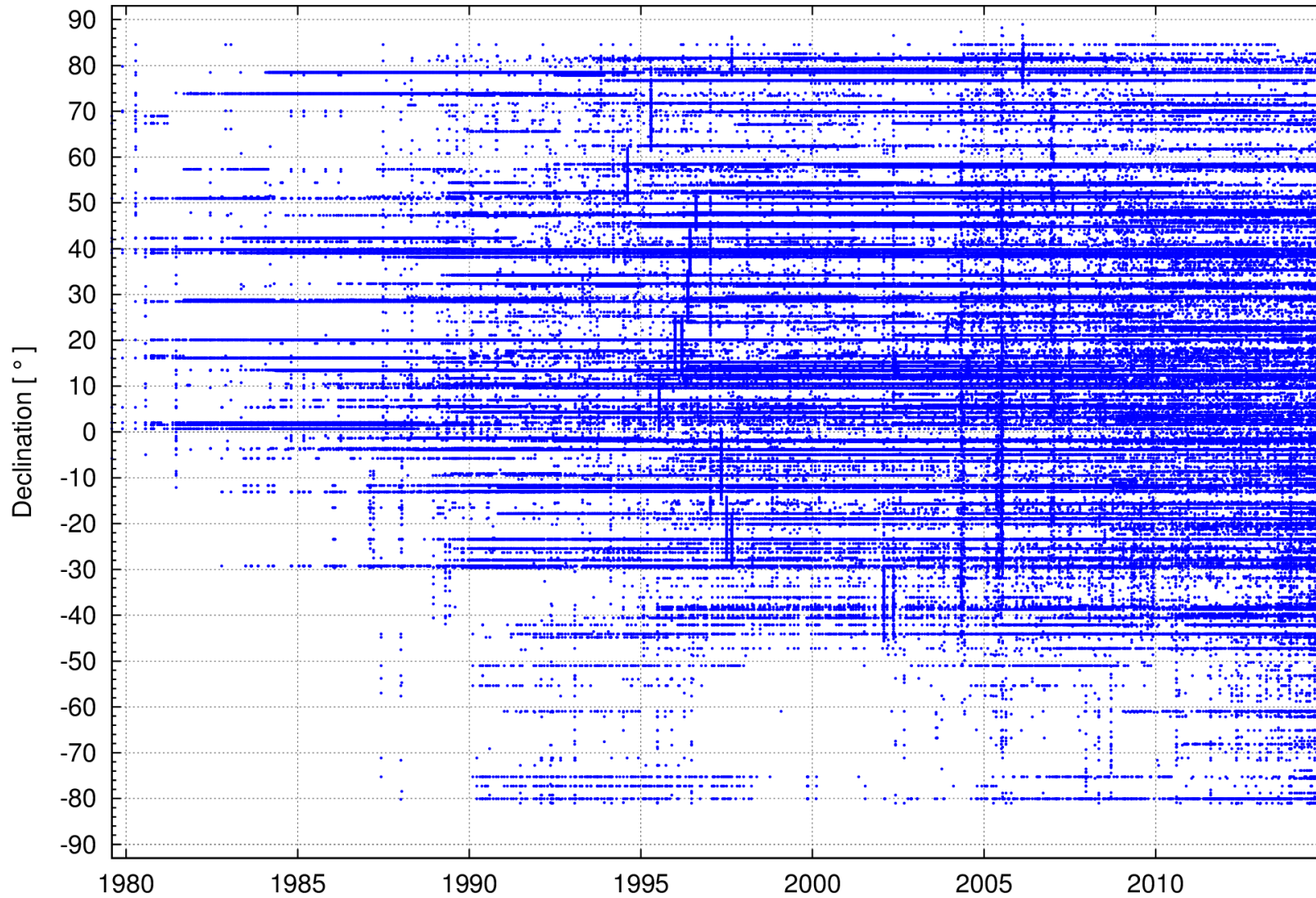
- From the immediate frame of the date (*having as pole the Earth rotation pole and possessing a non-rotating origin according to the celestial sphere*)
- To the celestial reference frame



⇒ **Precession-nutation**

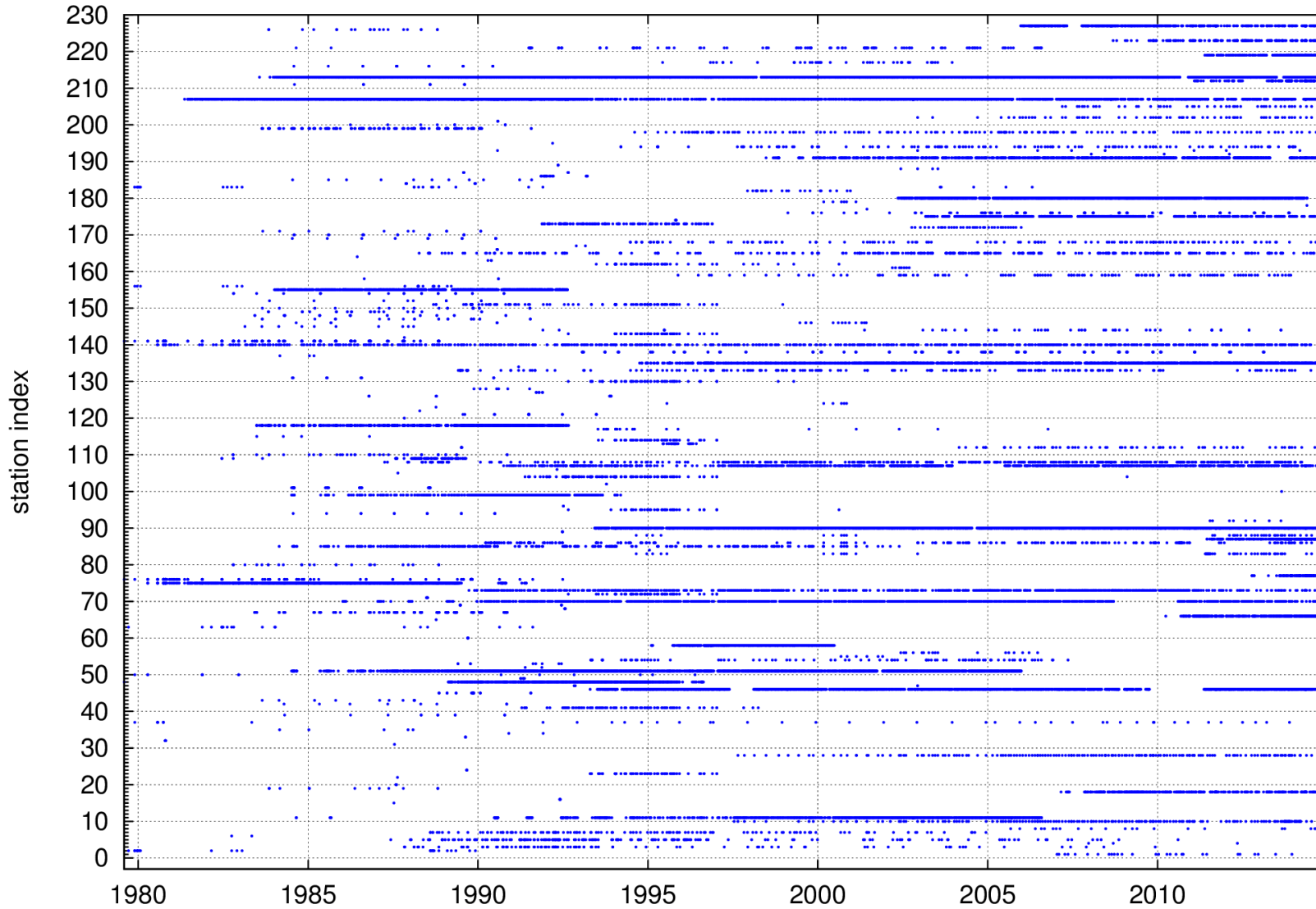
→ Parameters : **X** and **Y**

AVAILABLE SET OF DATA



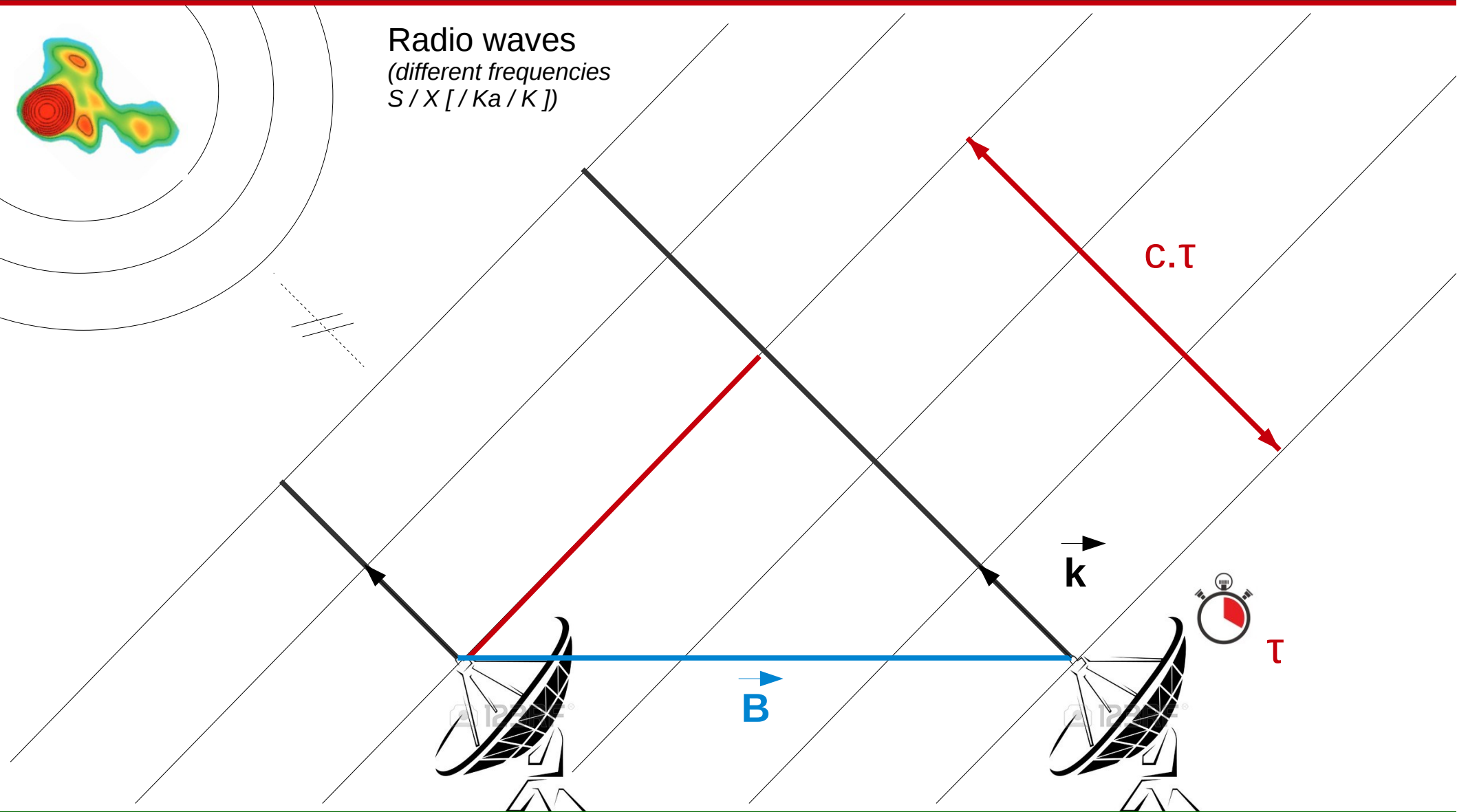
- **35** years of observation
(up to 24h duration)
- ~ **8000** VLBI sessions
- **5771** radiosources

AVAILABLE SET OF DATA

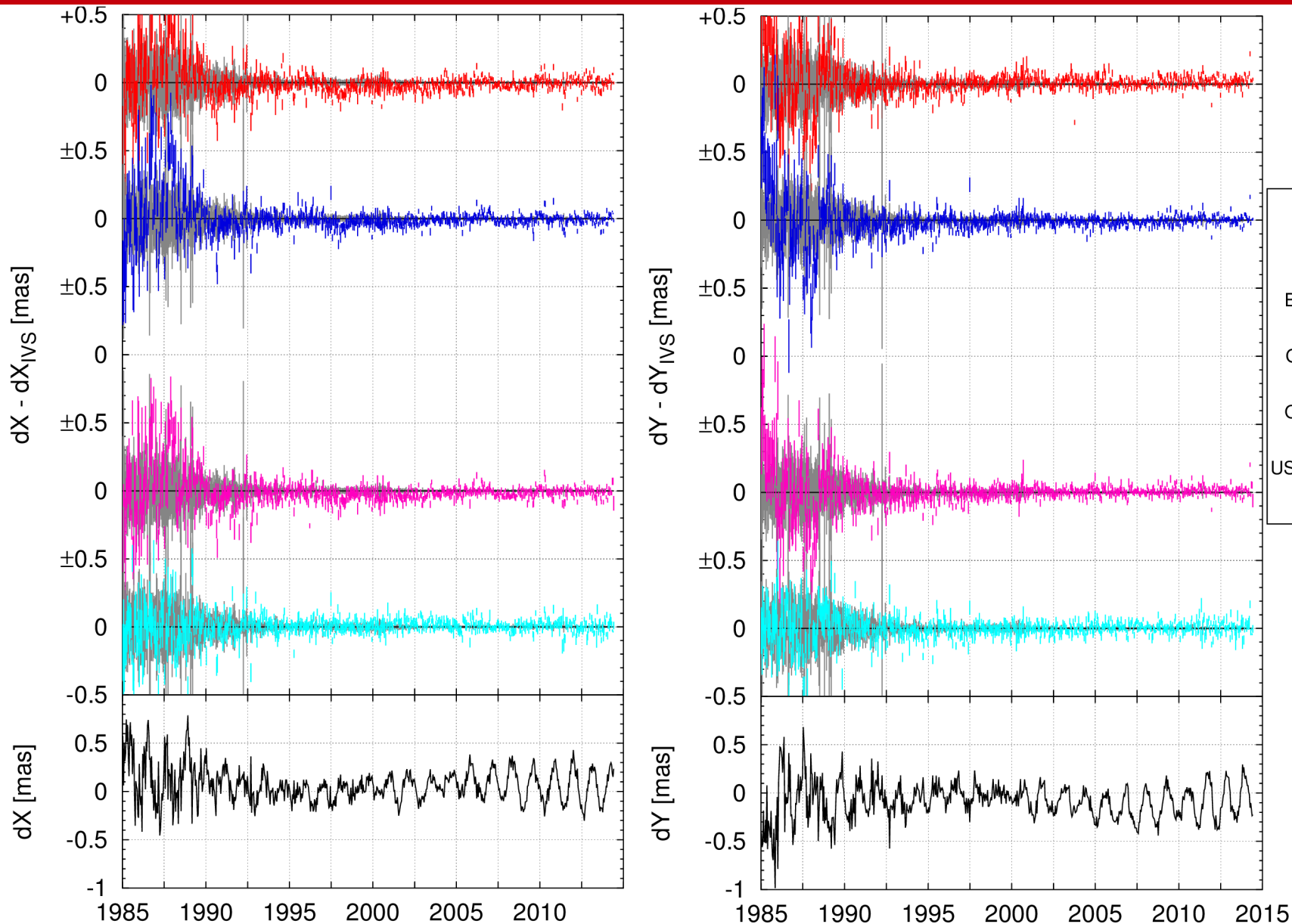


- **35** years of observation
(up to 24h duration)
- ~ **8000** VLBI sessions
- **227** stations
(fixed and mobile)

What is the VLBI principle ?



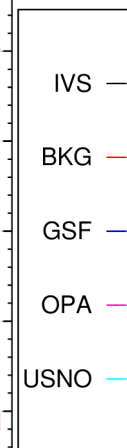
IERS Nutation Time Series - Comparison



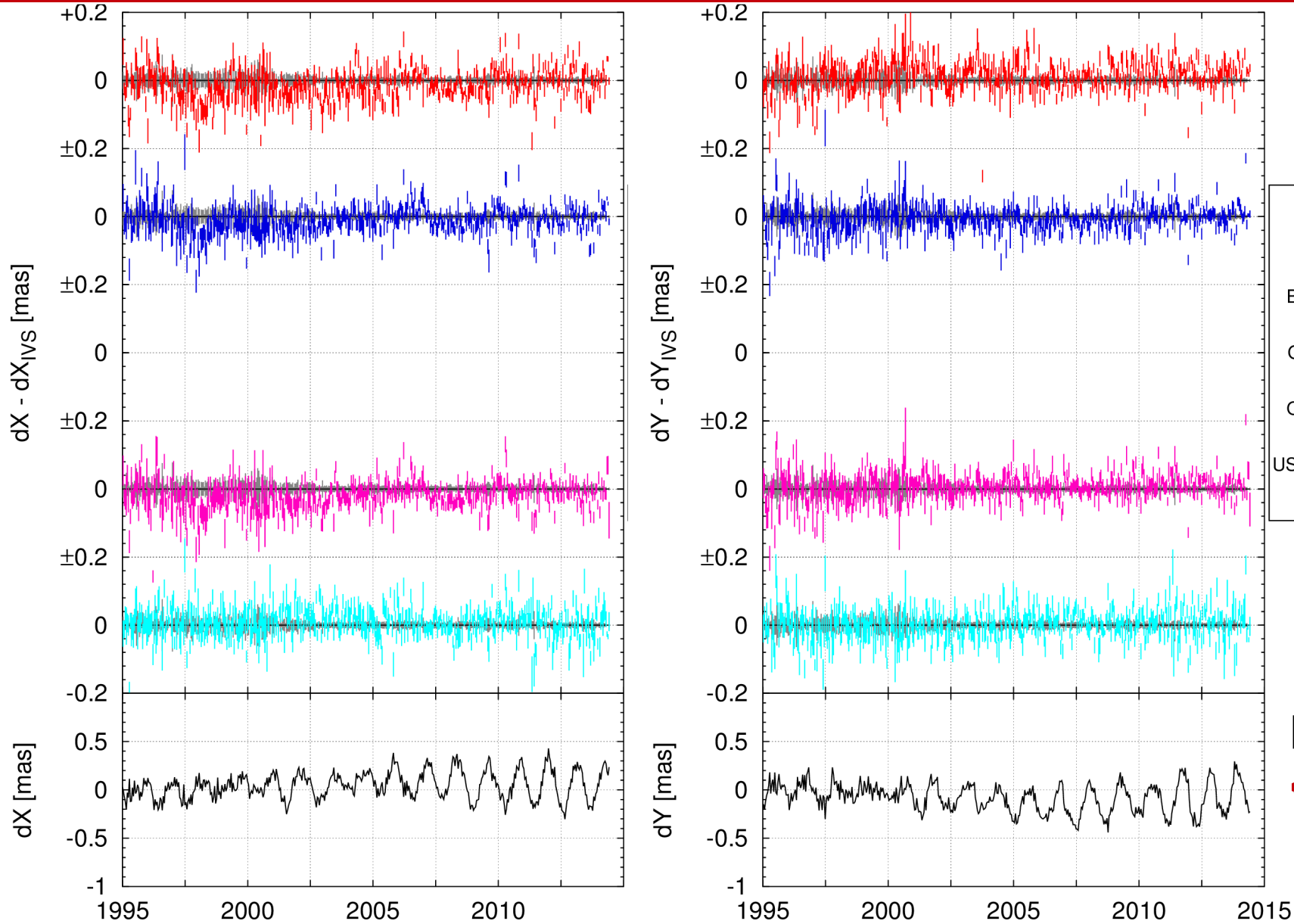
- Smoothed over 15 days

- errorbars for smoothed data are non inflated !

(some use 1.5 inflation)



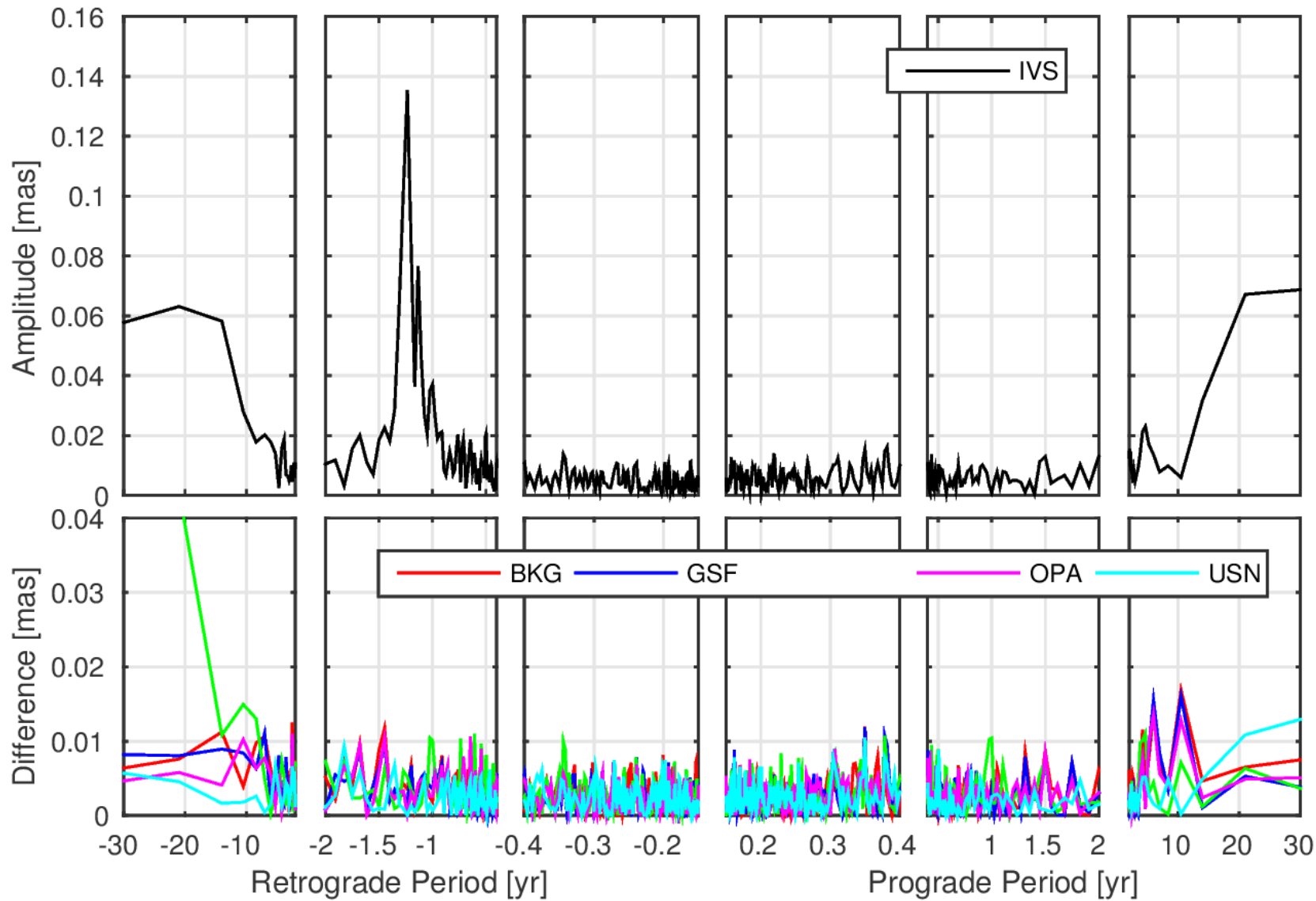
IERS Nutation Time Series - Comparison



- Smoothed over 15 days
- errorbars for smoothed data are non inflated !
- (some use 1.5 inflation)

Differences :
< 100 μ as

IERS Nutation Time Series - Spectra

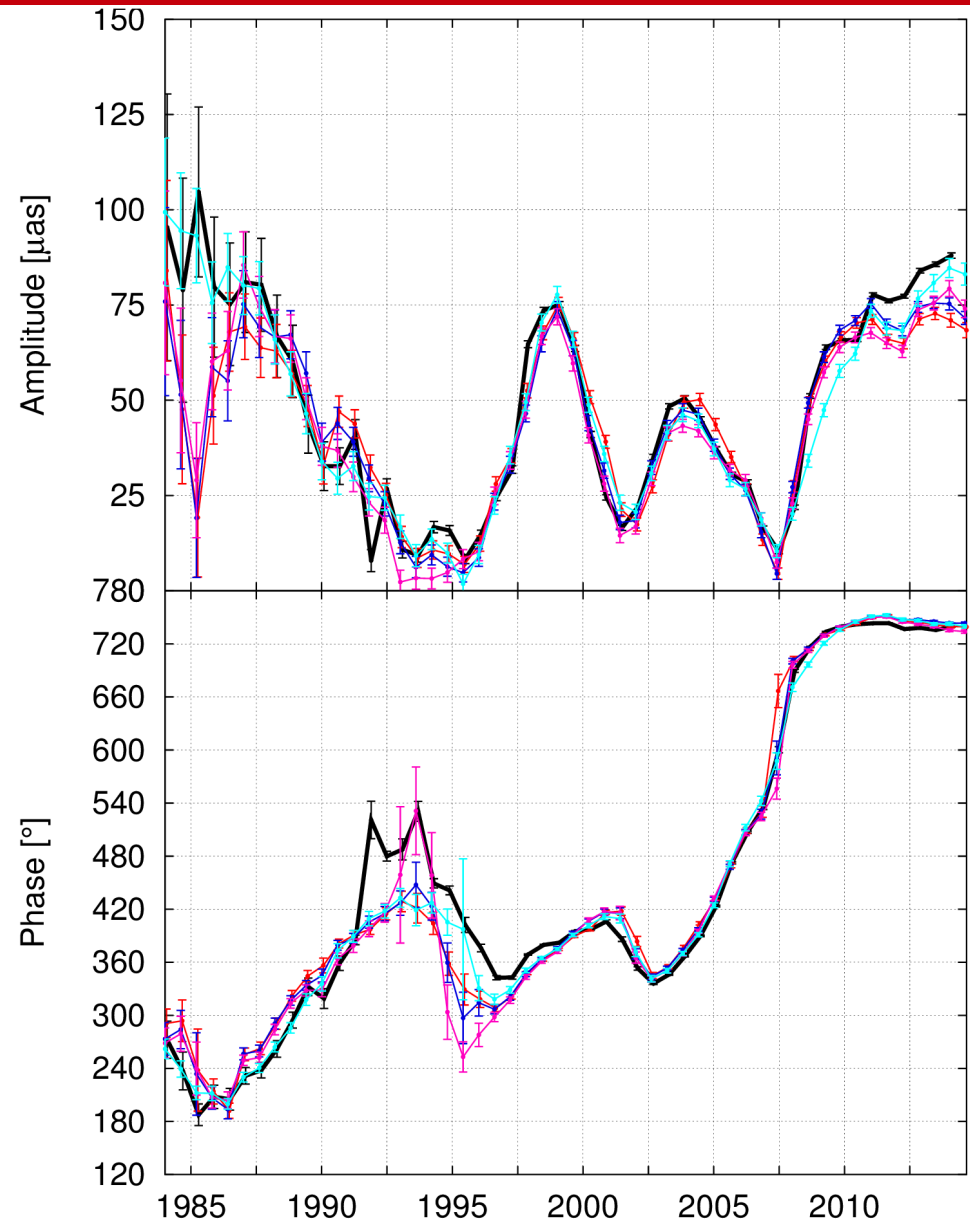
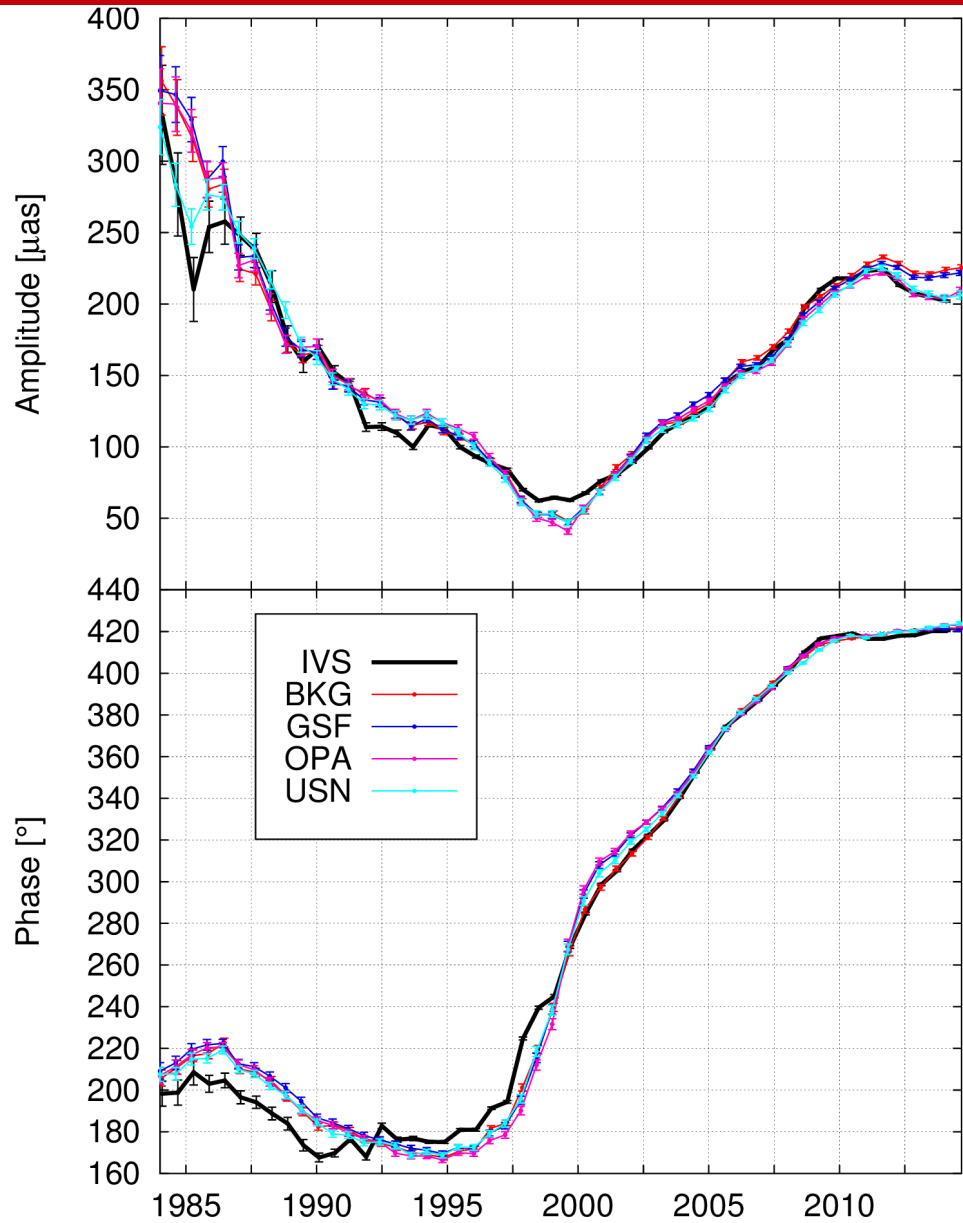


Amplitude differences :

< 10 μ as
for periods
< 2 years

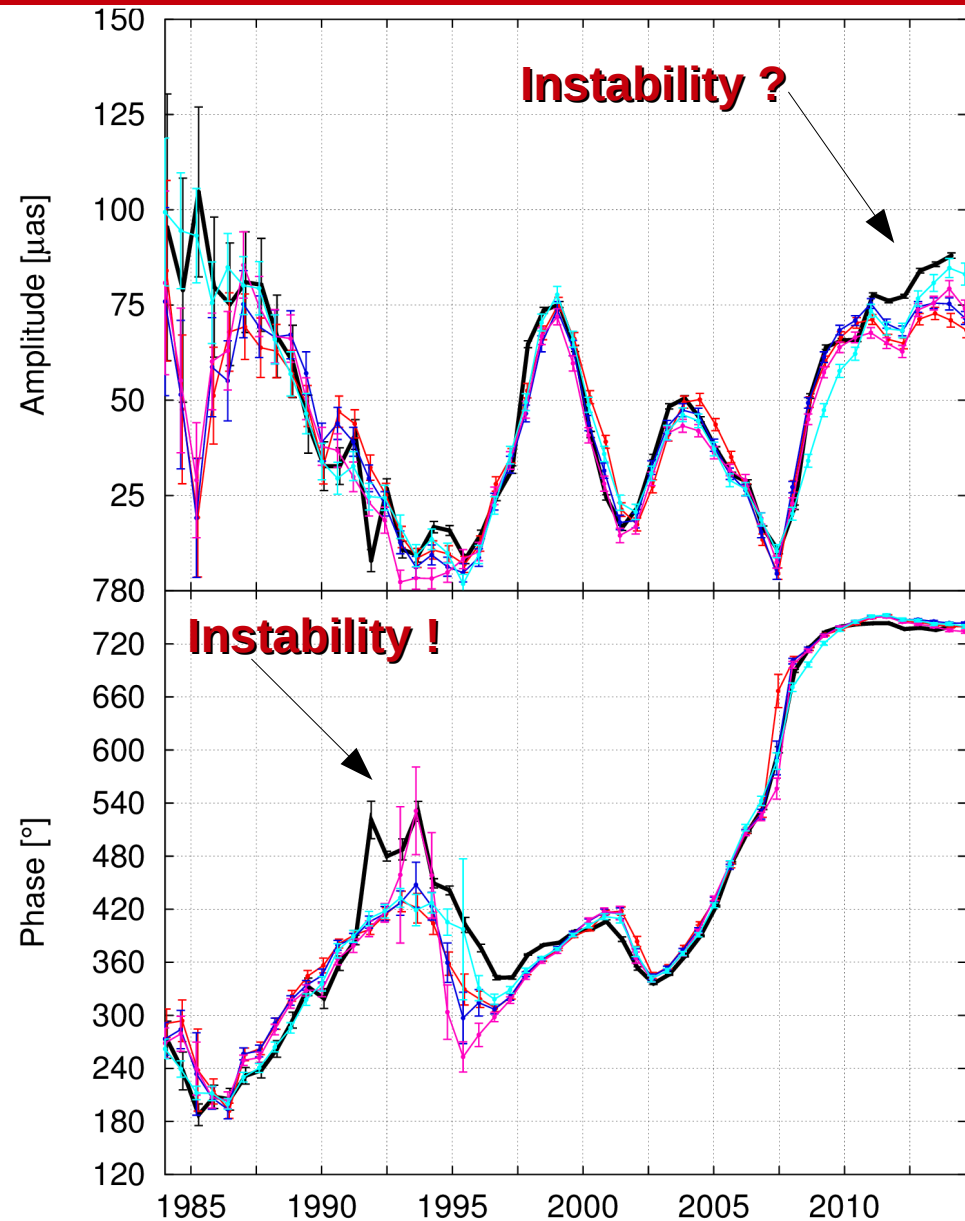
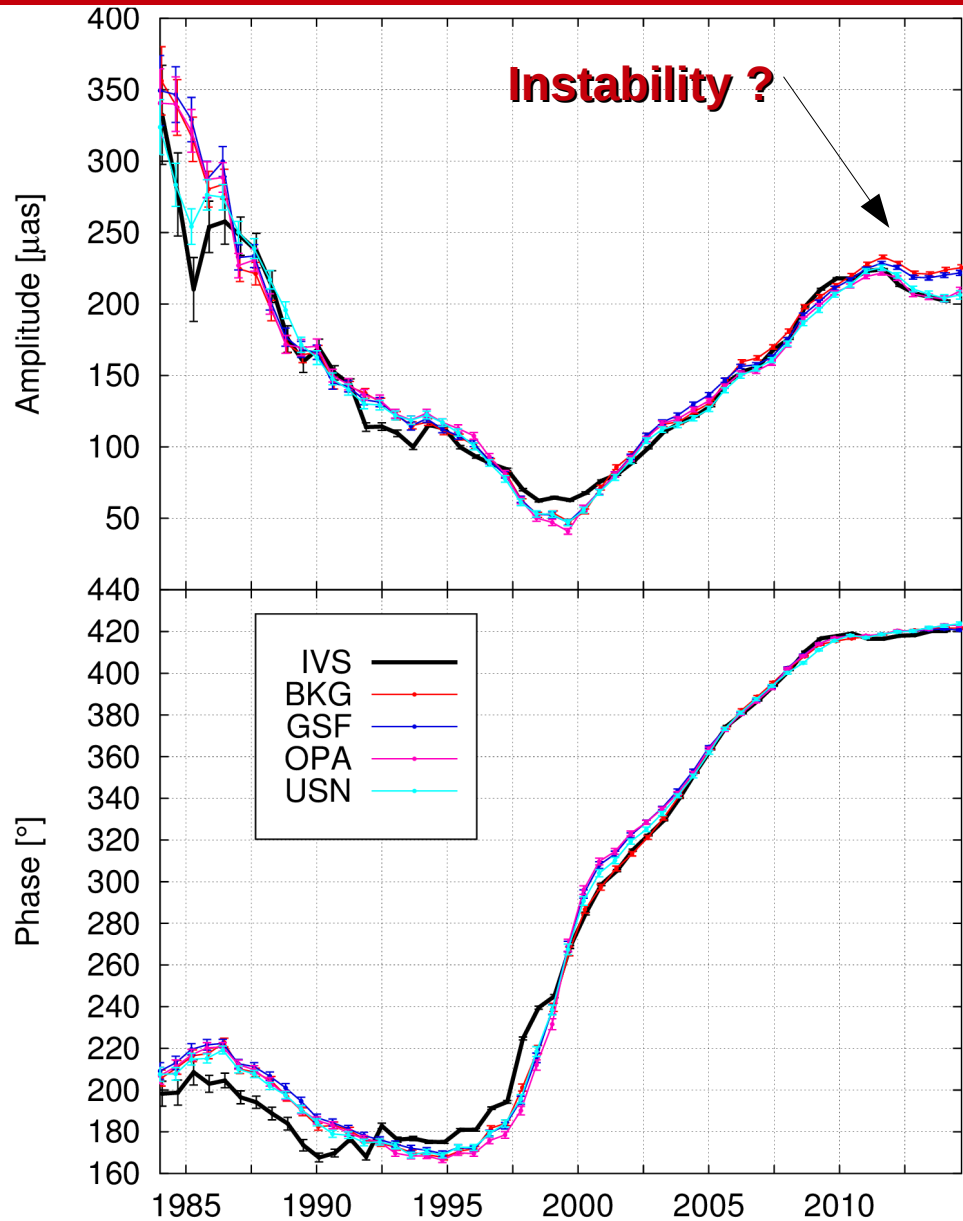
< 20 μ as
for periods
> 2 years

Free Core Nutation and Annual nutation Adjustments



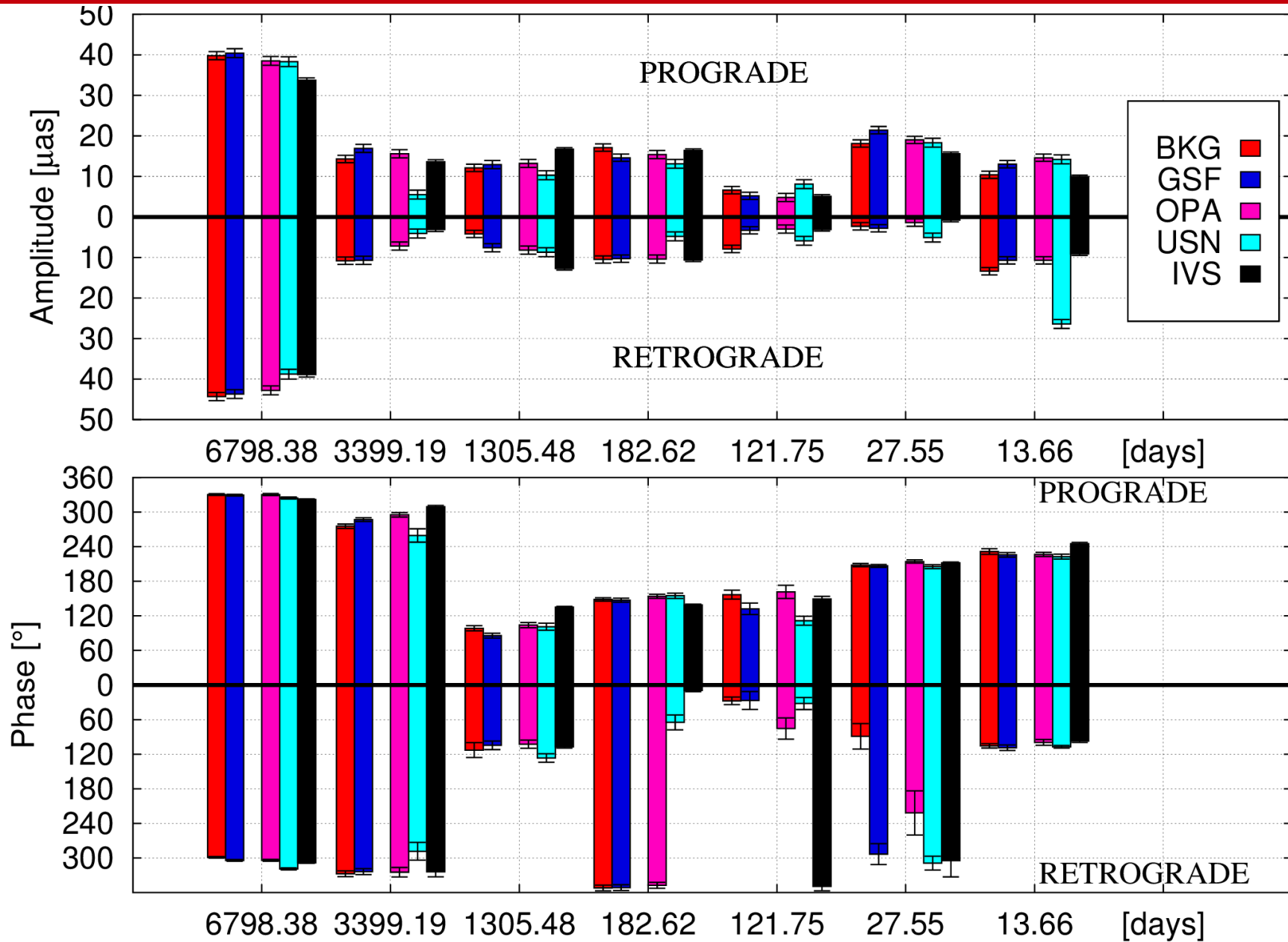
Stable over **10 μas** (amplitude), **10 $^\circ$** (FCN phase), **30 $^\circ$** (Annual Phase)

Free Core Nutation and Annual nutation Adjustments

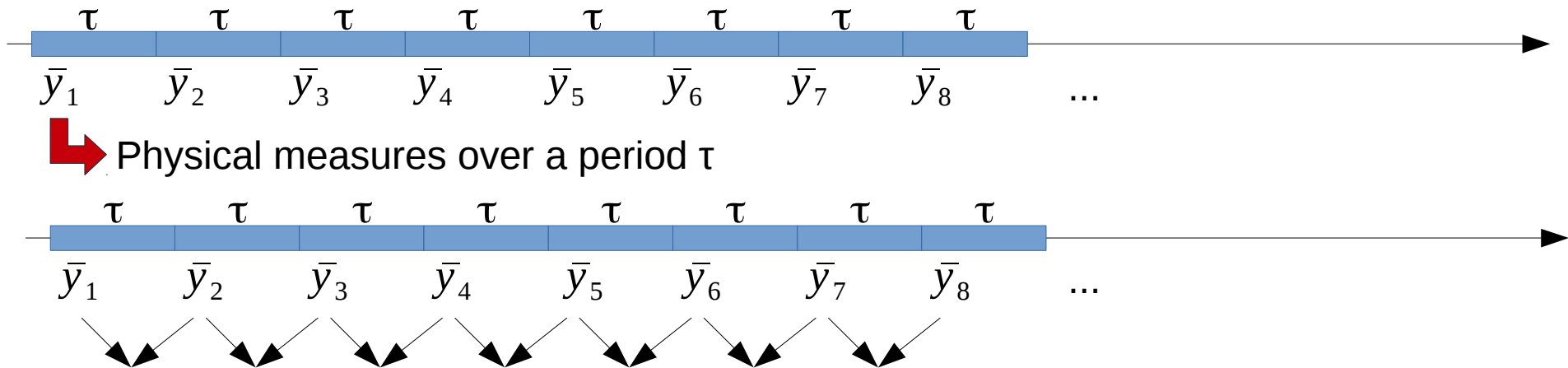


Stable over **10 μas** (amplitude), **10 $^\circ$** (FCN phase), **30 $^\circ$** (Annual Phase)

Others Principal Nutation Adjustements



IERS Nutation Time Series – Allan variance

Theoretical knowledge :

$$\frac{1}{2} (\bar{y}_k - \bar{y}_{k+1})^2 = \sigma_{(2)}^2(\tau) \quad \Rightarrow \quad \langle \sigma_{(2)}^2(\tau) \rangle$$

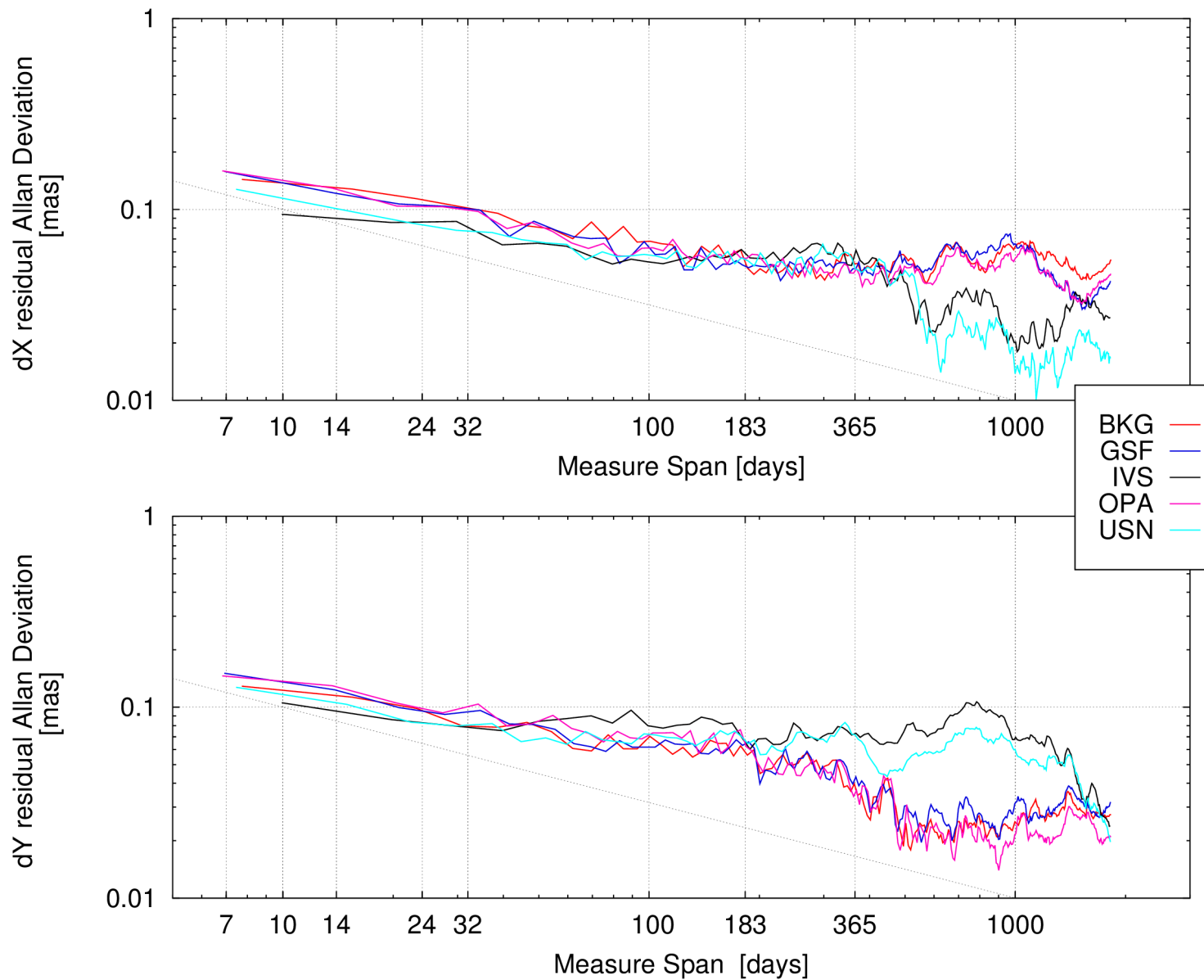
↳ Allan Variance measurement

↳ Estimation of the true variance of the data set given τ

The Allan variance serie with respect to τ allow to :

- determine **the type of noise** associated with datas
- estimate **the true variance** of datas

IERS Nutation Time Series – Allan Deviation

**Result :****Flicker Noise**

+

periodic signal at long period (> 5 years)

Or

White Noise

+


Outliers

Remark : A rejection of outliers has been done beforehand

Center Analysis Strategy

	BKG	GSF	OPA	USN
CRF	a priori ICRF2 NNR 295 Def. sources ?? global / ?? local	A priori gsf2012a.src NNR 295 Def. sources 1670 global / 39 local	A priori ICRF2 NNR 295 Def. sources ?? global / 39 local	A priori ICRF2 NNR 295 Def. sources 846 global / 852 local
Nutation	A priori IAU2006/2000A Apply recommendation IERS Convention 2010	A priori IAU2006/2000A Apply recommendation IERS Convention 2010	A priori IAU2006/2000A Apply recommendation IERS Convention 2010	A priori IAU2006/2000A Apply recommendation IERS Convention 2010
Tropo	Zenith : 1h linear spline VMF1 wet partial derivative (segmented) A priori made by make_vmf_trp_file from GSFC based on VMF1 Gradient : East and north offset A priori from DAO model	Zenith : 20-min linear spline VMF wet partial derivative (segmented) A priori VMF total mapping function Saastamoinen model Gradient : 6-hour linear spline east and north A priori from DAO model	Zenith : 20-min linear spline A priori VMF1 mapping function Gradient : 6-hour east and north offset A priori from DAO model	Zenith : 20-min linear spline NMF wet partial derivative (segmented) A priori NMF dry mapping function Saastamoinen model Gradient : 6-hour linear spline at all station except 110 A priori from DAO model
Clock	1h linear spline	Quadratic (local) + 1h linear spline (segmented)	Quadratic (local) + 1h linear spline	Quadratic (local) + 1h linear spline
Elevation cutoff	5° elevation cutoff	5° elevation cutoff	5° elevation cutoff	5° elevation cutoff
Software	CALC 11.01 Solve 2014.02.21	CALC 11 SOLVE 2014.02.21	CALC 11.0 SOLVE 2014.02.21	CALC 11 SOLVE 2014.02.21

Conclusion

- There exist differences between nutation time series of IERS **at the order of 100 μs**
- Those are consequences of differences in nutation adjustments **at the order of 10 μs** in amplitude and **at the order of 10-30°** in phase
- Residuals after adjustments seem to be animated by a **flicker noise** with a periodic signal at long period (> 5 years)

- Signal are stable **at 100 μs over 7 days** and at **few tenth of μs over period of several years**

Thanks

**I thank you for
your attention**

If you want to contact me :

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